Eugene F. Milone Denis A. Leahy David W. Hobill Editors

Short-Period Binary Stars: Observations, Analyses, and Results





WJ Hussar

Short-Period Binary Stars: Observations, Analyses, and Results Eugene F. Milone, Denis A. Leahy, David W. Hobill, 2008-02-21 Short period binaries run the gamut from widely separated stars to black hole pairs in between are systems that include neutron stars and white dwarfs and partially evolved systems such as tidally distorted and over contact systems These objects represent stages of evolution of binary stars and their degrees of separation provide critical clues to how their evolutionary paths differ from that of single stars The widest and least distorted systems provide astronomers with the essential precise data needed to study all stars mass and radius The interactions of binary star components on the other hand provide a natural laboratory to observe how the matter in these stars behaves under different and often varying physical conditions Thus cataclysmic variables with and without overpoweringly strong magnetic fields and stars with densities from that found in the Sun to the degenerate matter of white dwarfs and the ultra compact states of neutron stars and black holes are all discussed The extensive index permits cross referencing Eclipsing Binary Stars: Modeling and Analysis Josef Kallrath, Eugene F. Milone, 2009-08-14 Have you ever stopped at a construction project on the way to your of ce and the day's astrophysics Remember the other onlookers folks just enjoying the sp tacle as we all do in following developments away from our areas of active work We are excited and thrilled when the Hubble Space Telescope discovers an Einstein Cross when the marvelous pulsars enter our lives and when computer scientists put a little box on our desk that outperforms yesterday s giant machines We are free to make use of such achievements and we respect the imagination and discipline needed to bring them about just as onlookers respect the abilities and planning needed to create a building they may later use After all each of us contributes in our own areas as best as we can In addition to the serious onlookers there will be passersby who take only a casual look at the site They may use the building later but have little or no interest in its construction and give no thought to the resources needed tobring it to completion Upon arriving at work those persons write astronomy and astrophysics books at various levels in which they must say something about close binary stars Usually a page or two will do and the emphasis is on the MLR mass luminosity radius data obtained only from binaries Short-Period Binary Stars: Observations, Analyses, and Results Eugene F. Milone, Denis A. Leahy, David W. Hobill, 2009-09-03 This book explores cataclysmic variables with and without strong overpowering magnetic fields You ll read about stars with densities ranging from that of the Sun to the degenerate matter of white dwarfs to the ultra compact states of neutron stars and black holes One of the objects examined and discussed is the Double Pulsar highlighting what observations have told us INIS Atomindex ,1984 about fundamental physics Physics Briefs ,1988 Research Centers Directory, 1979 Research institutes foundations centers bureaus laboratories experiment stations and other similar nonprofit facilities organizations and activities in the United States and Canada Entry gives identifying and descriptive information of staff and work Institutional research centers and subject indexes 5th ed 5491 entries 6th ed 6268 entries Scientific and

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periodic brightness variations which he found Over this long span of time our knowledge about these systems has been growing and we have now reached a fairly good understanding of the structure and behavior of this interesting group of objects This book contains a timely summary of our present knowledge of interacting binary stars The chapters have been written by distinguished scientists who have done relevant research in the field of interacting binary stars Modeling of Eclipsing Binary Stars E.F. Milone, 2012-12-06 In the two decades since the development of the first eclipsing binary modeling code new analytic techniques and the availability of powerful sometimes dedicated computing facilities have made possible vastly improved determinations of fundamental and even transient stellar parameters. The scale of these developments of course raises questions about modeling tools techniques and philosophies such as Who will maintain and upgrade the codes Will the codes be open to improvement by outsiders and if so how And indeed what should be the goals of a modeling program Such questions had not been aired for a long time and for this reason alone deserved to be discussed in as general a forum as the community provides This volume contains material presented by Commission 42 Close Binary Stars during the International Astronomical Union s XXI General Assembly in Argentina July 1991 and during IAU Colloquium 151 Cordoba Argentina August 1991 The techniques discussed include simulations of stellar bright and dark spots streams partial and complete stellar disks prominences and other features characterizing active stars modeling of polarization parameters models that use radial velocities as well as line profile simulations to model velocity field variation across stellar disks the weighted effects of brightness asymmetries and models for translucent eclipsing agents such as stellar winds

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